

# LMT Performance

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A large satellite dish antenna is shown from a low angle, looking up against a dark blue sky. The dish is metallic and has a complex support structure. The text 'Outline' is written in white on the left side of the dish.

# Outline

- Antenna Gain
- Pointing
- Summary

# Antenna Gain

*some definitions*

- Geometric Area :  $A_G = \frac{\pi}{4} D^2 = 1963.5 \text{ m}^2$

- Effective Area:  $A_E = \frac{2kT_A}{S}$

- S is flux density of calibration source
- $T_A$  is antenna temperature observed on source
- Effective area is always less than geometric area due to illumination of aperture by receiver, antenna surface errors, blockage and shadowing of the aperture.

- Aperture Efficiency:  $\eta_A = \frac{A_E}{A_G}$

- Sensitivity Factor:  $G = \frac{2k}{A_E} Jy K^{-1}$

- Beam Efficiency (Main)  $\eta_B = \frac{\iint_{MB} B(x,y) dx dy}{\iint_{-\infty}^{+\infty} B(x,y) dx dy}$

$B(x,y)$  is the power pattern of the antenna



# Antenna Gain

*Expected Frequency  
Dependence*

- Frequency Dependence of Effective Area is well described by *Ruze Theory*<sup>1</sup>

$$A_E(\nu) = A_E(0) e^{-\left(\frac{4\pi\nu}{c}\epsilon\right)^2}$$

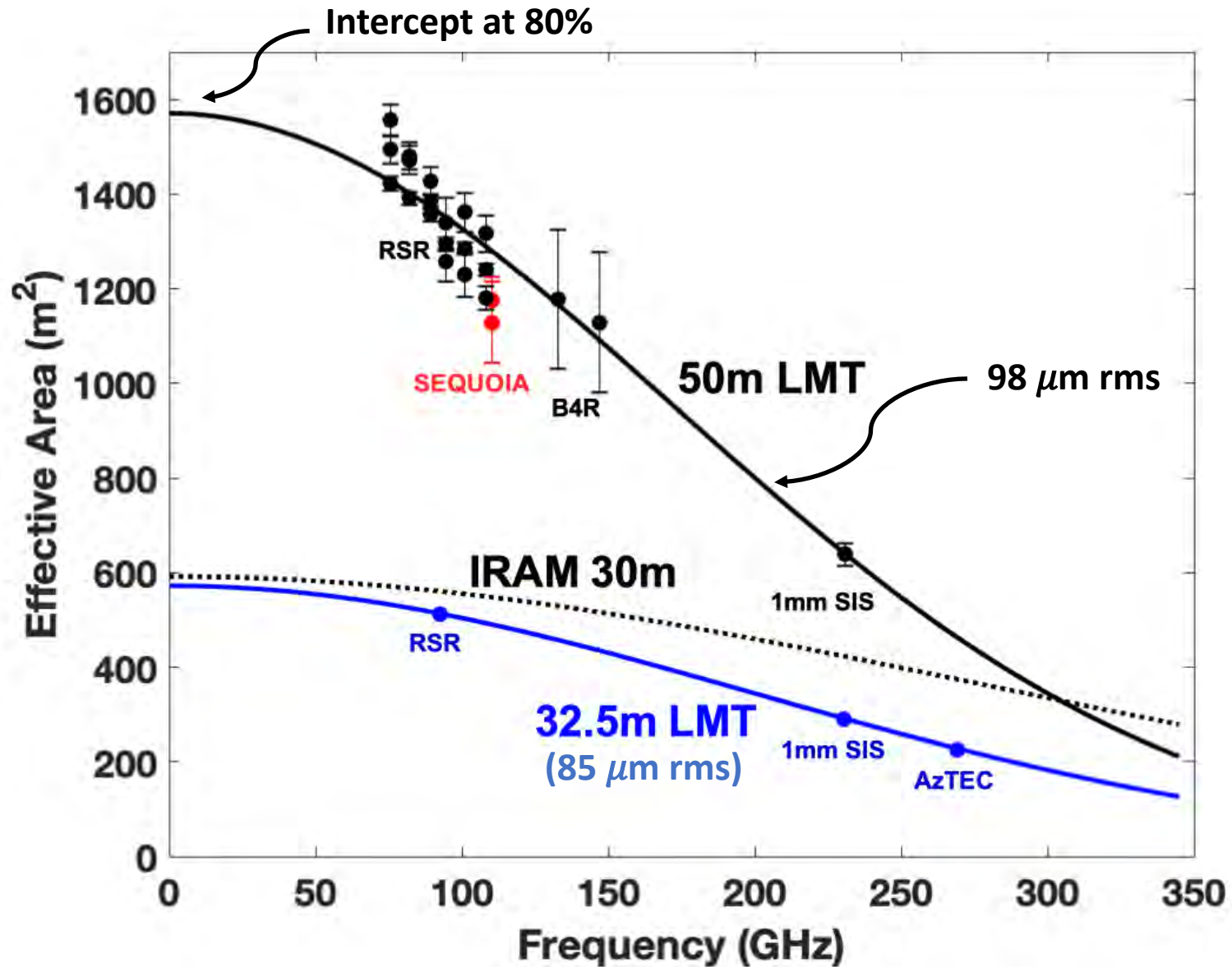
where  $\epsilon$  is rms surface error;  $\nu$  is frequency

- $A_E(0)$  is effective area at 0 frequency. It includes effects of:
  - antenna blockage and shadowing
  - receiver illumination of primary

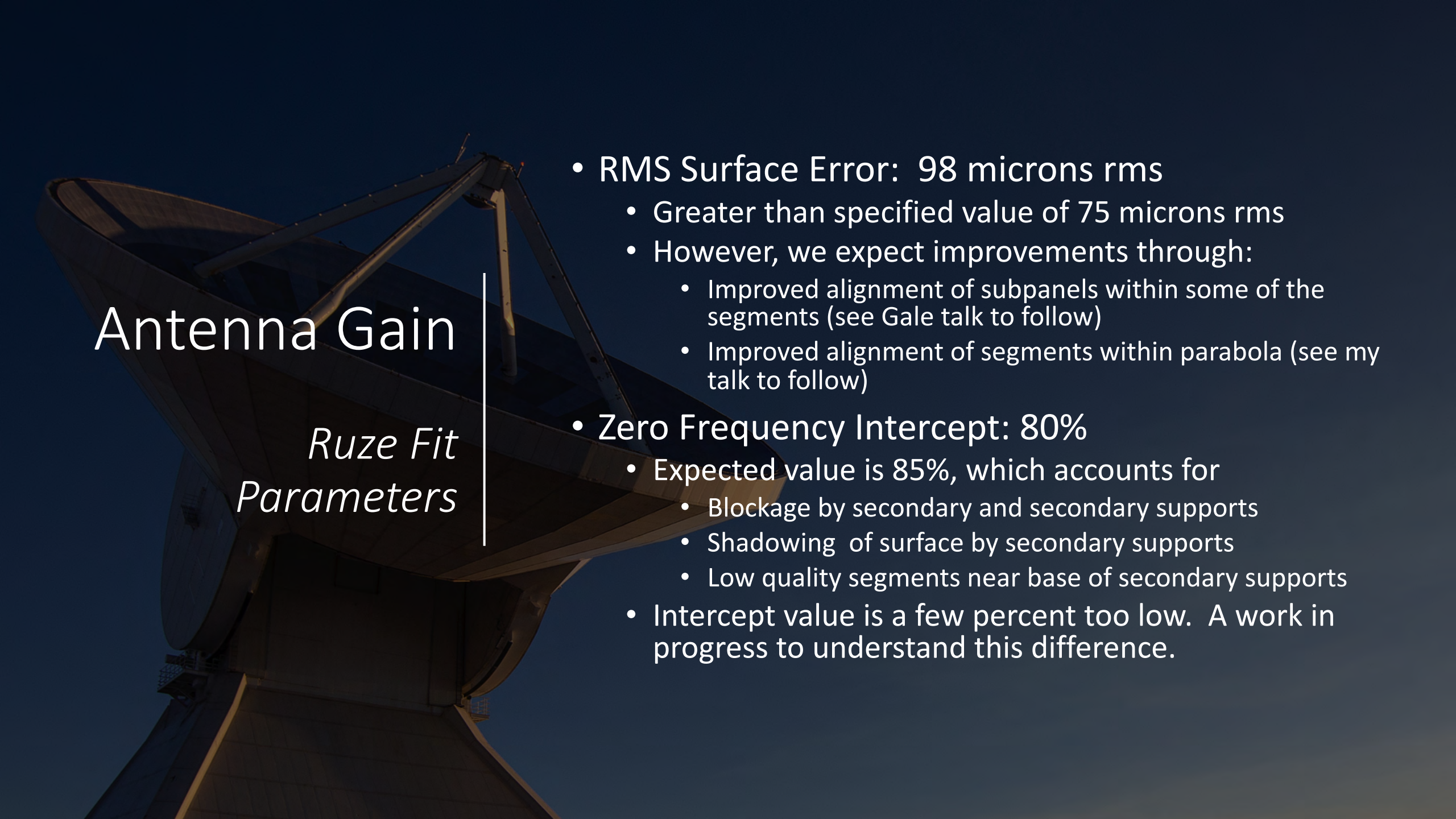
<sup>1</sup> J. Ruze, Proceedings of the IEEE, vol. 54, pp 633-642 (1966)

# Antenna Gain

## Ruze Plot



Note: Effective area is corrected for receiver illumination of primary in this plot.



# Antenna Gain

## *Ruze Fit Parameters*

- RMS Surface Error: 98 microns rms
  - Greater than specified value of 75 microns rms
  - However, we expect improvements through:
    - Improved alignment of subpanels within some of the segments (see Gale talk to follow)
    - Improved alignment of segments within parabola (see my talk to follow)
- Zero Frequency Intercept: 80%
  - Expected value is 85%, which accounts for
    - Blockage by secondary and secondary supports
    - Shadowing of surface by secondary supports
    - Low quality segments near base of secondary supports
  - Intercept value is a few percent too low. A work in progress to understand this difference.



# Antenna Gain

*Aperture and  
Main Beam  
Efficiency*

<b>FREQUENCY (GHz)</b>	$\eta_A$	$\eta_{MB}$	<b>G (Jy/K)</b>
85	0.58	0.68	2.42
100	0.55	0.65	2.56
115	0.52	0.61	2.70
140	0.47	0.55	2.99
230	0.27	0.31	5.21

Aperture efficiency based on Ruze fit; Main beam efficiency is estimated



## Absolute Pointing

Ability to point at any location in the sky

We think this is about 2" rms

## Antenna Pointing

*some definitions*

## Relative (Offset) Pointing

Ability to point after a pointing check on a nearby (~10 degrees) radio source

We think this is about 1" rms

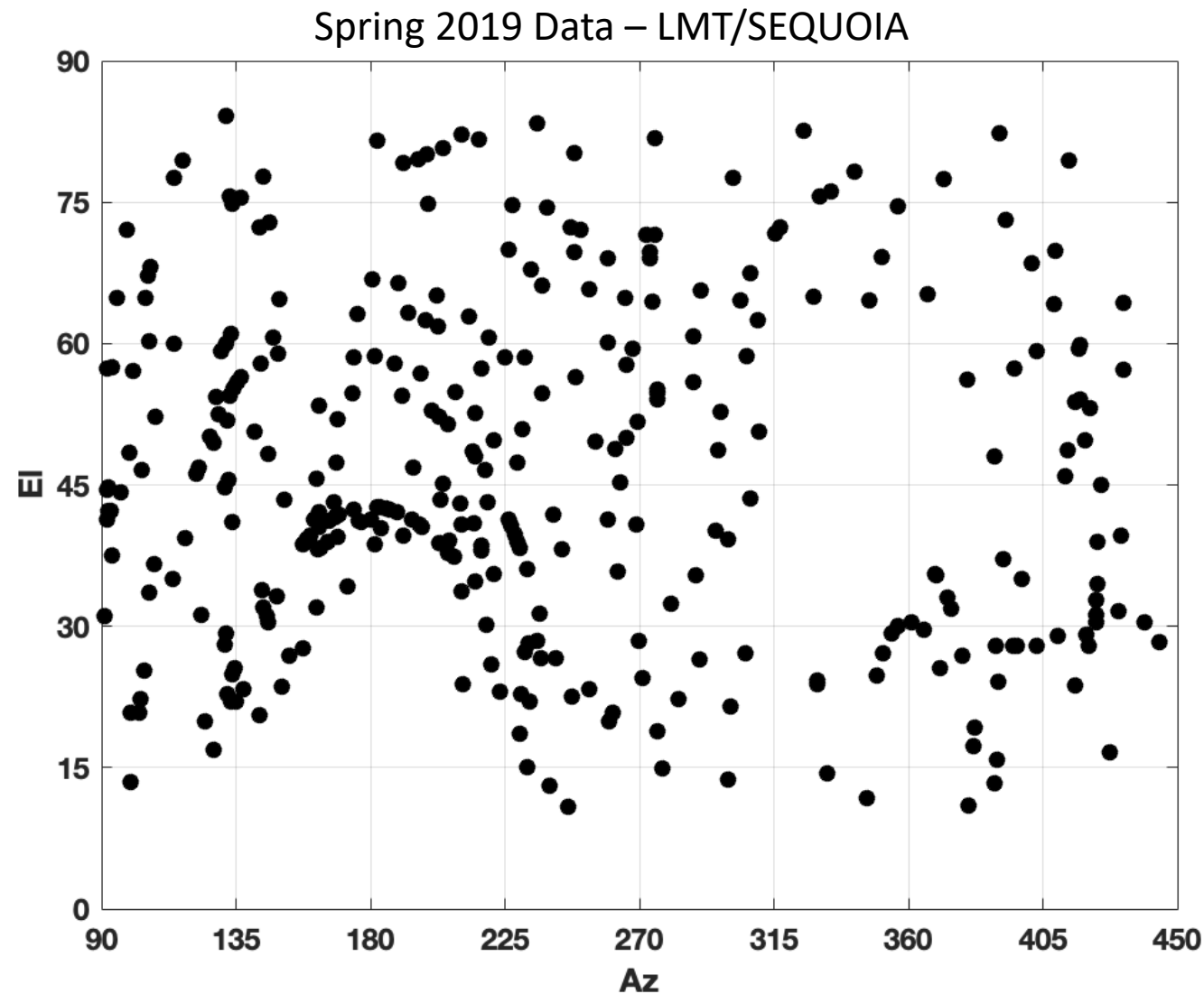
Note: Pointing is not strongly affected by winds less than operational design limit of 10 m/s.



# Antenna Pointing

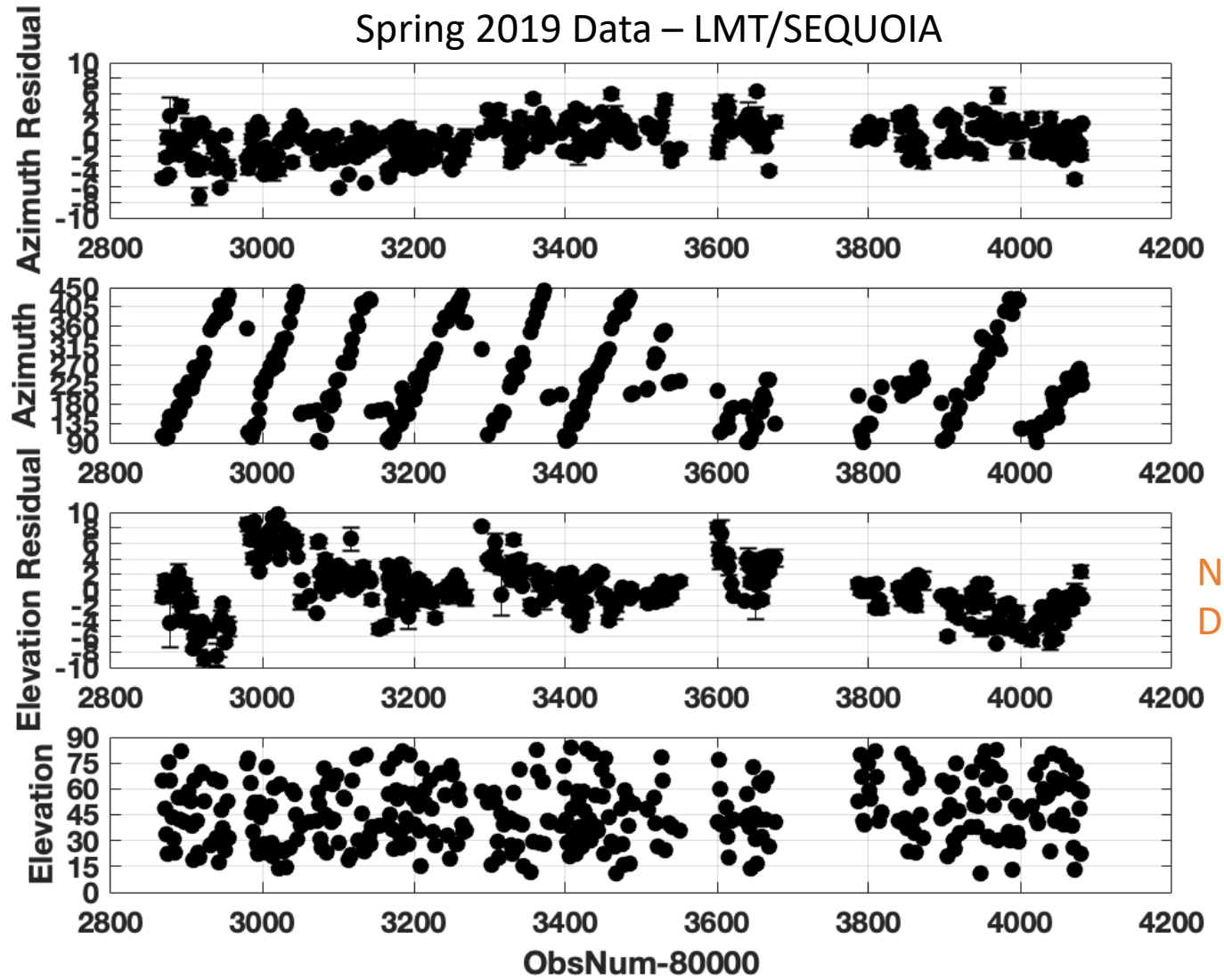
*All Sky Data*

*Spring 2019*  
*5 nights*  
*LMT SEQUOIA*



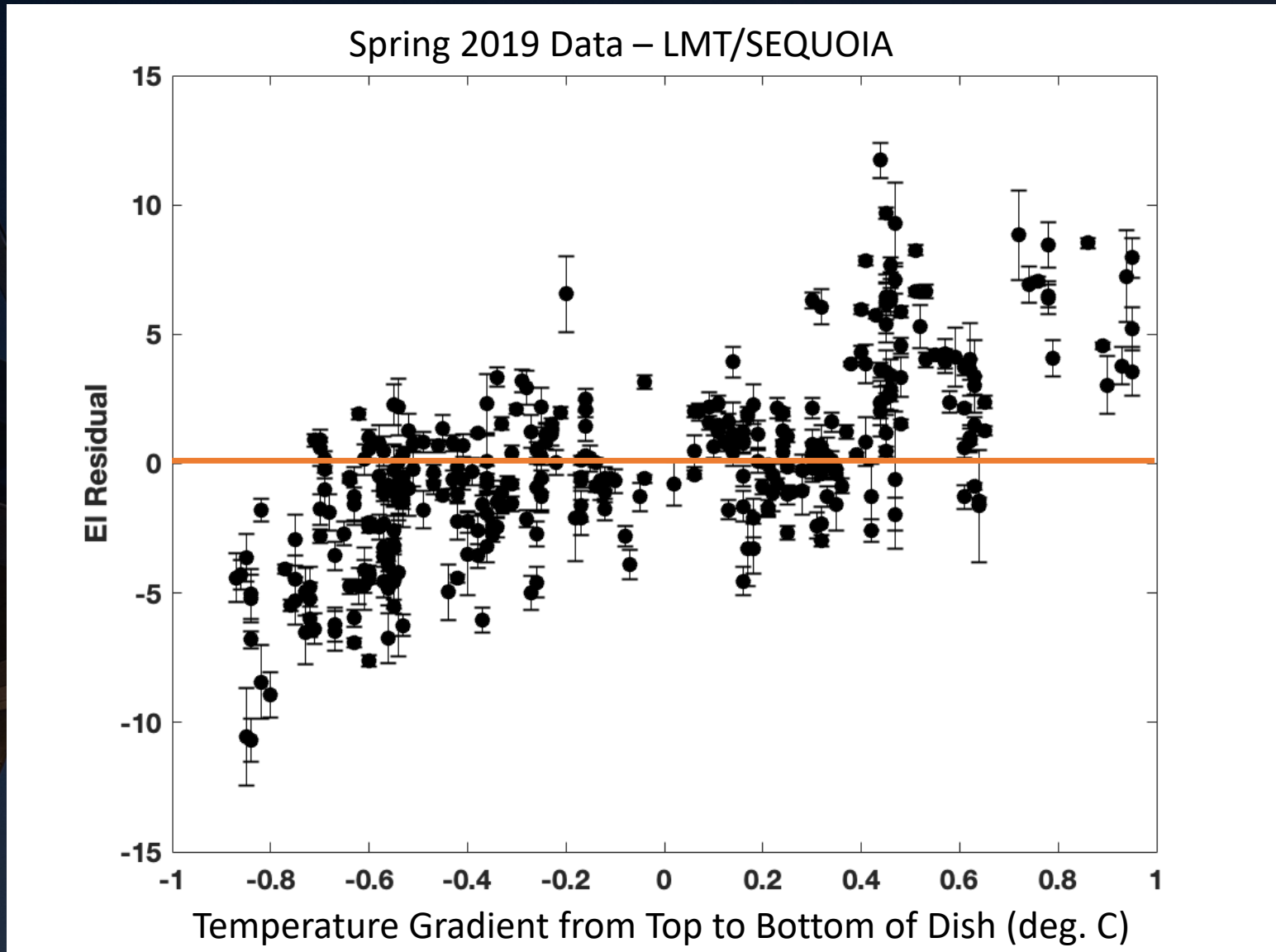
# Antenna Pointing

*Residuals to All Sky Pointing Model*



# Antenna Pointing

*Residuals compared to temperature gradients*



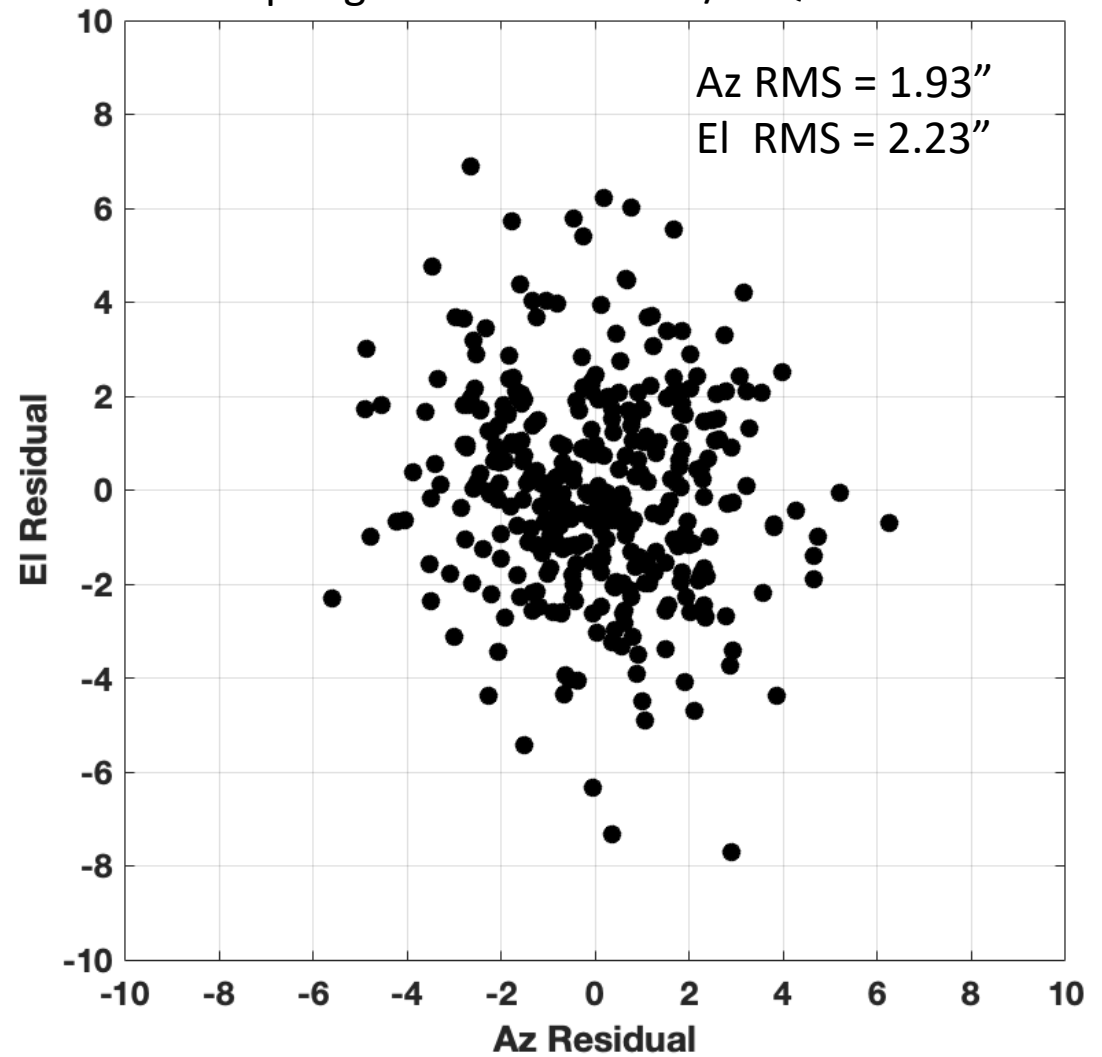
*Observed drifts are probably thermal in nature*

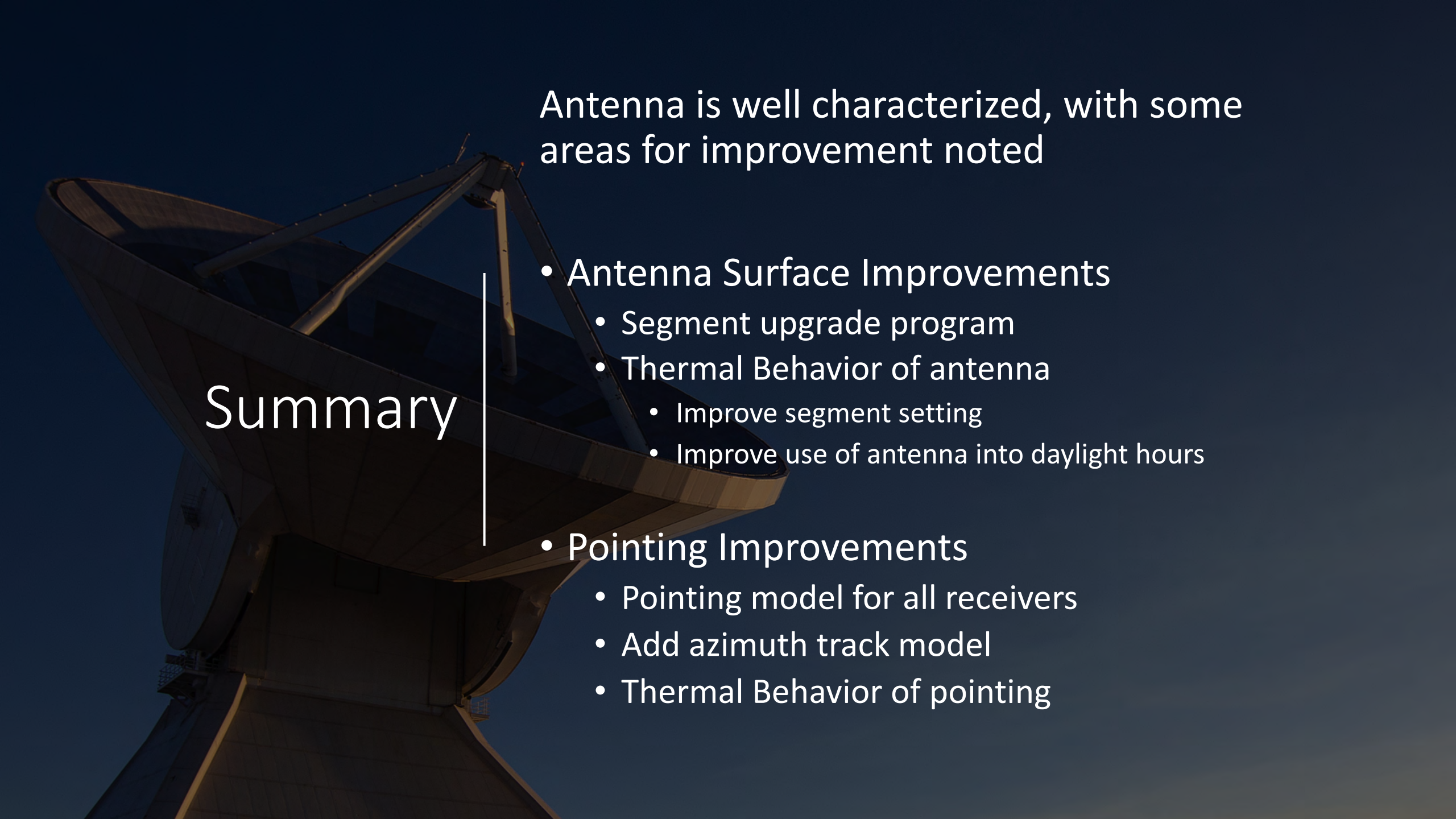
# Antenna Pointing

*Pointing Residuals*

*(assume repointing periodically during night)*

Spring 2019 Data – LMT/SEQUOIA





Antenna is well characterized, with some areas for improvement noted

## Summary

- Antenna Surface Improvements
  - Segment upgrade program
  - Thermal Behavior of antenna
    - Improve segment setting
    - Improve use of antenna into daylight hours
- Pointing Improvements
  - Pointing model for all receivers
  - Add azimuth track model
  - Thermal Behavior of pointing